

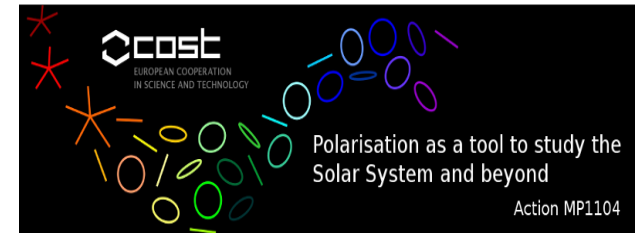
The perspectives of X-ray polarimetry on accreting sources



(with MoCA: a Monte Carlo code for Comptonization in Astrophysics)



FP7
Strong Gravity Project
WP4 “Comptonization models:
spectra and polarization”



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Stefano Bianchi²

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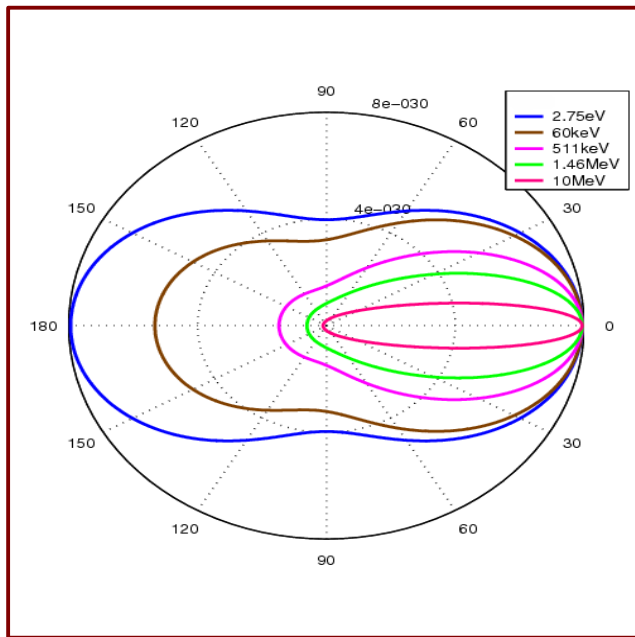
² University of Rome “Roma Tre”

³ Astronomical Institute of the Academy of Science (Prague)

Outline

- Scientific Motivation
- The model
- The continuum emission in AGN
- The Iron line broadening in NS
- Conclusions & observational perspectives

Scientific Motivation



Why another code?

Except for a few cases

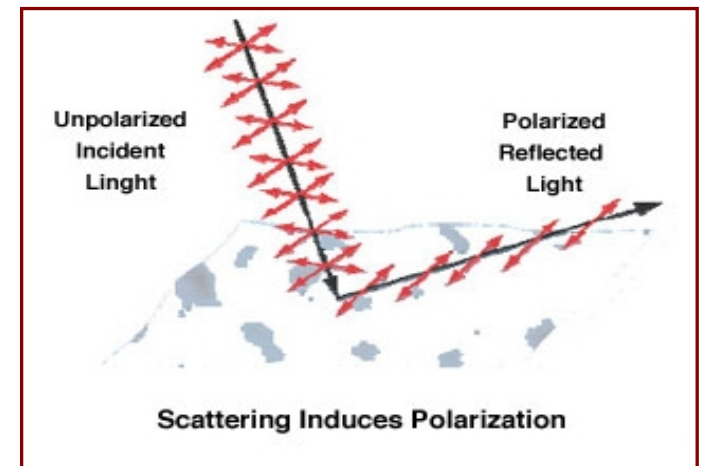
- no Klein-Nishina cross-section
- no polarization

MoCA includes both

Why polarimetry?

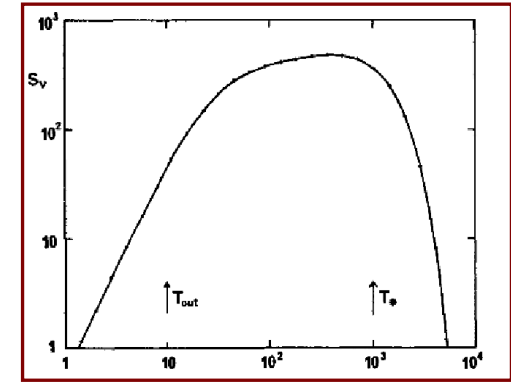
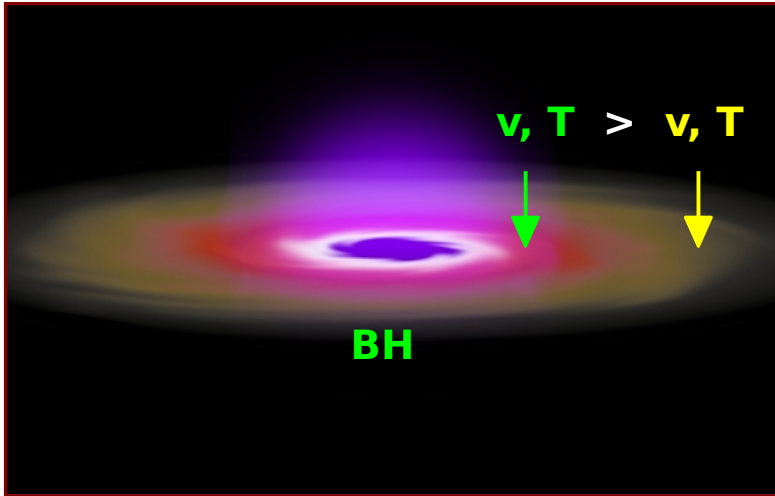
Observational astronomy is mainly made through EM radiation using 3 observables: **wavelength**, **time** and **space** (i.e. spectroscopy, timing and imaging)

polarimetry adds **2 more** independent observables which are very sensitive to **geometry** and **inclination**



Comptonization

Multi-temperature BB emission



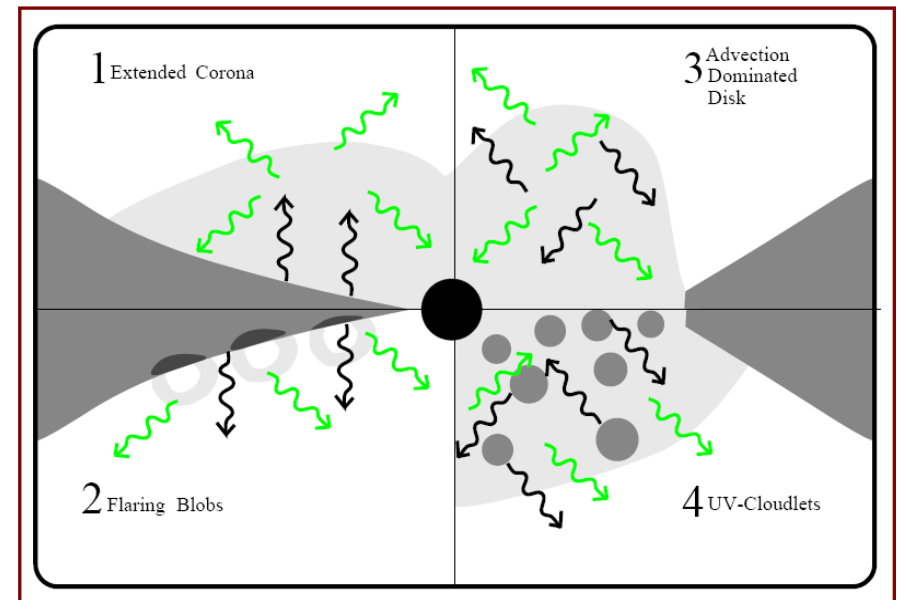
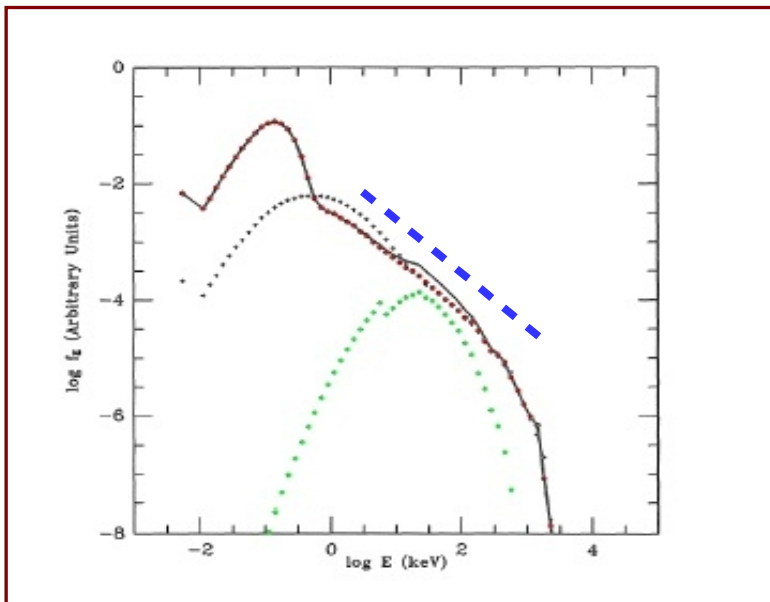
Galactic BHs ($10 M_{\odot}$) \rightarrow soft X-rays

SMBHs ($10^8 M_{\odot}$) \rightarrow UV



Steady, geometrically thin & optically thick disc
(Shakura-Sunyaev 1973)

Comptonization by relativistic
(thermal) electrons

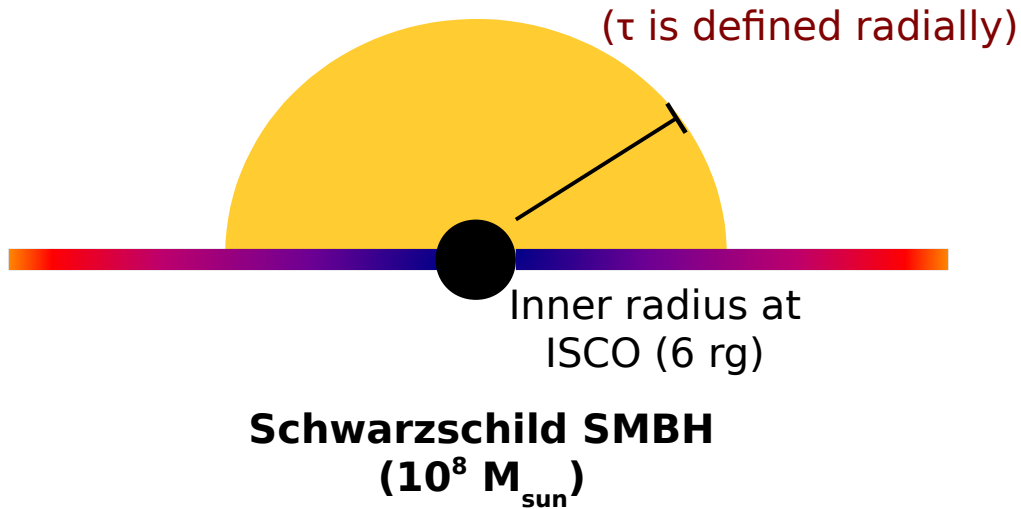


Two-phase disc (Haardt-Maraschi 1991)

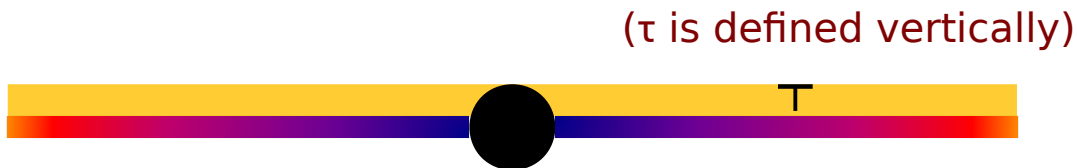
The parameters of the corona are
basically unknown !

The Model

HEMISPHERICAL corona

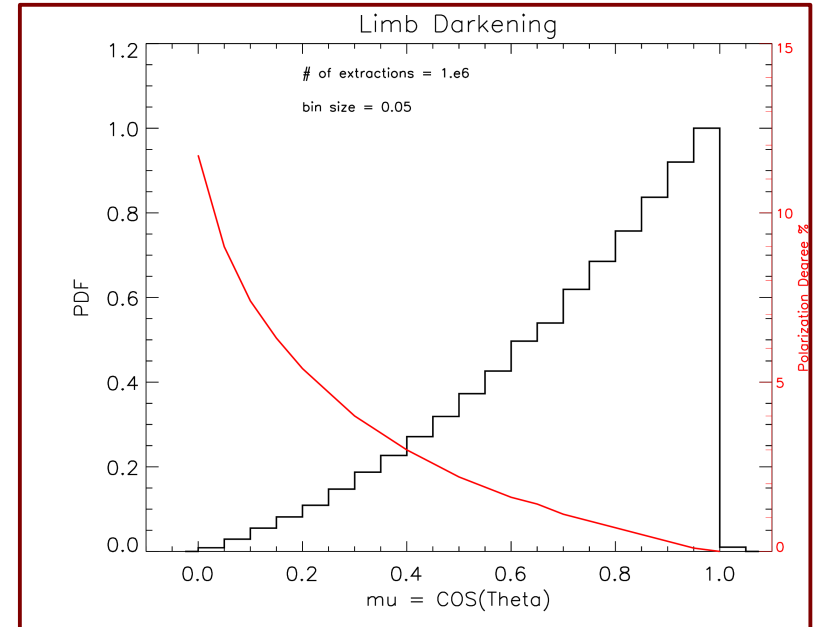


SLABBY corona



Seed photons:

- limb darkening
- **horizontal (i.e. 0°) polarization**



(Chandrasekhar 1960)

Corona parameters:

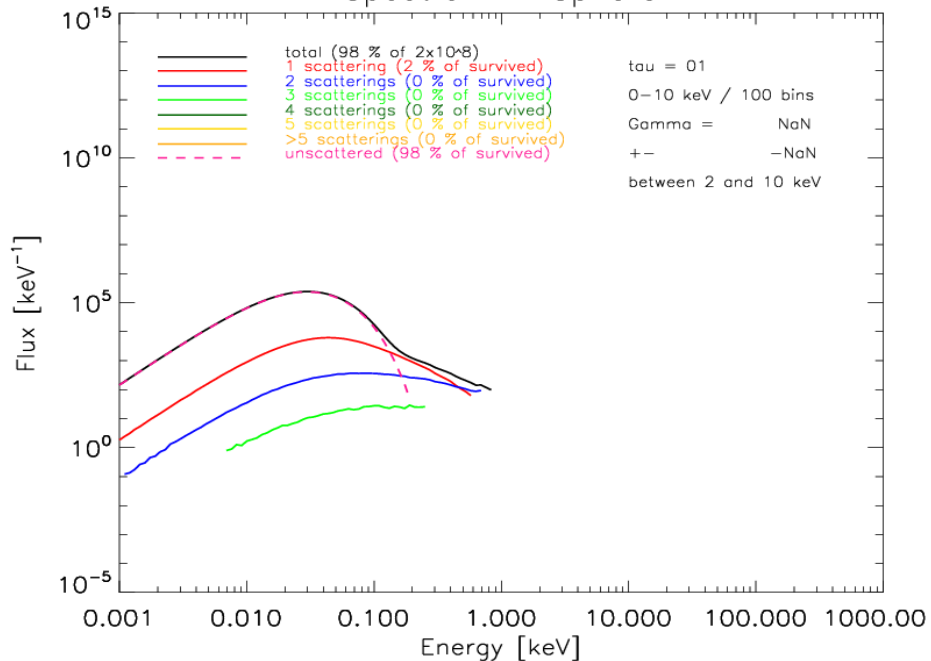
- thermal energy ($kT = 100 \text{ keV}$)
- optical depth ($\tau = \mathbf{0.1, 1}$)

SPHERE

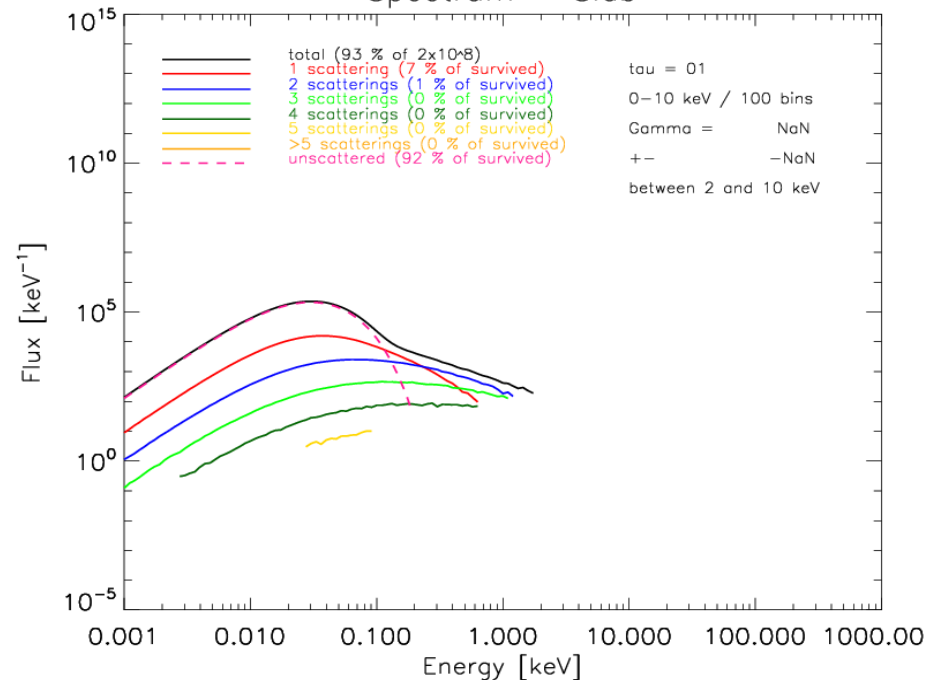
AGN Spectra

SLAB

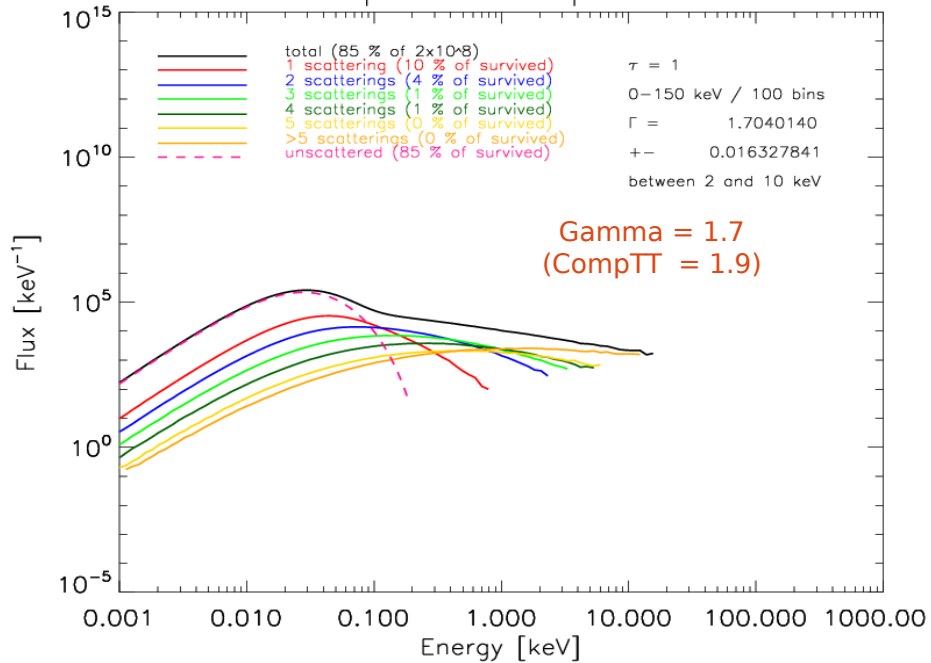
Spectrum – Sphere



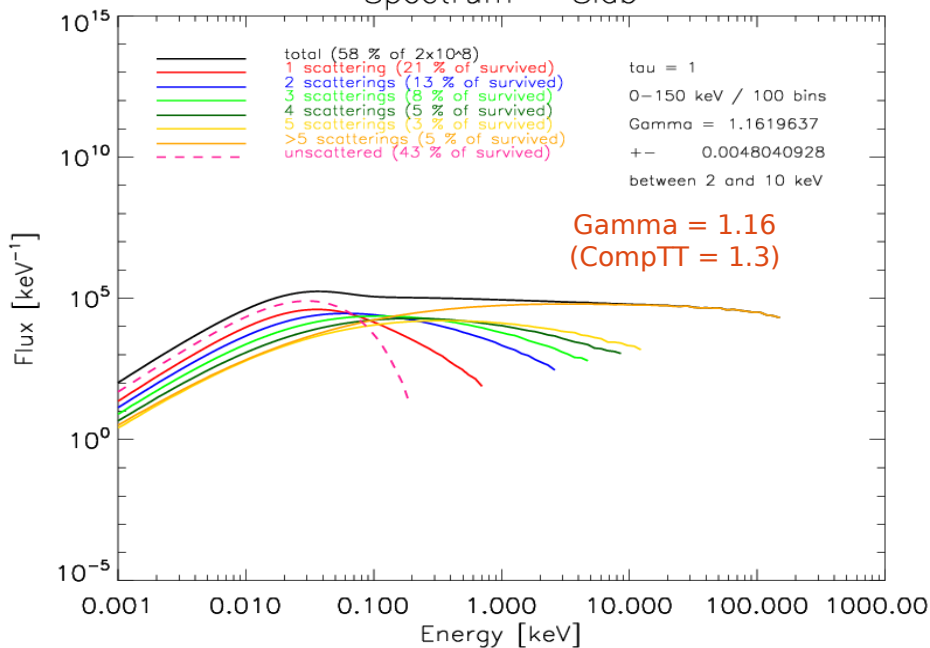
Spectrum – Slab



Spectrum – Sphere



Spectrum – Slab



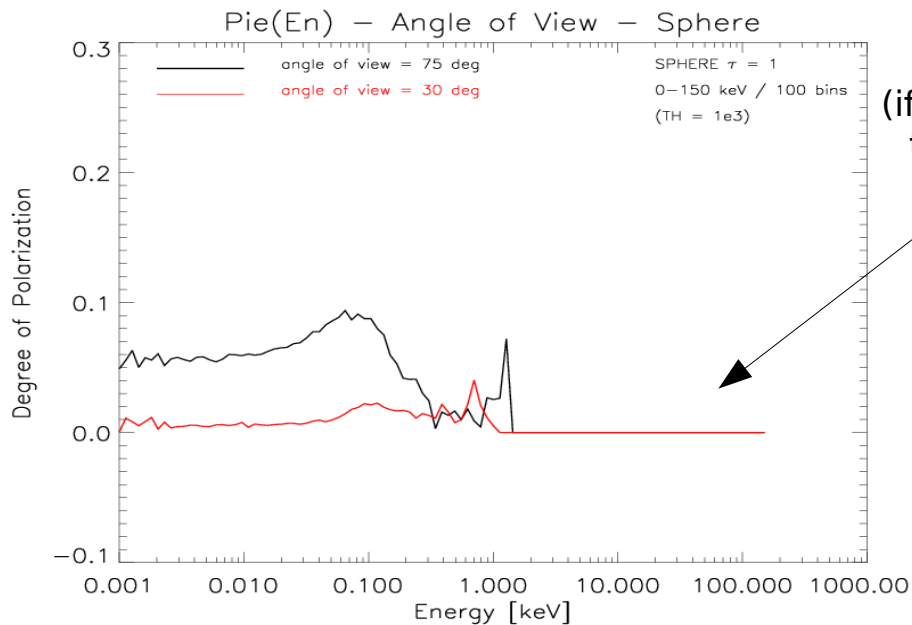
SPHERE

$\tau = 1$

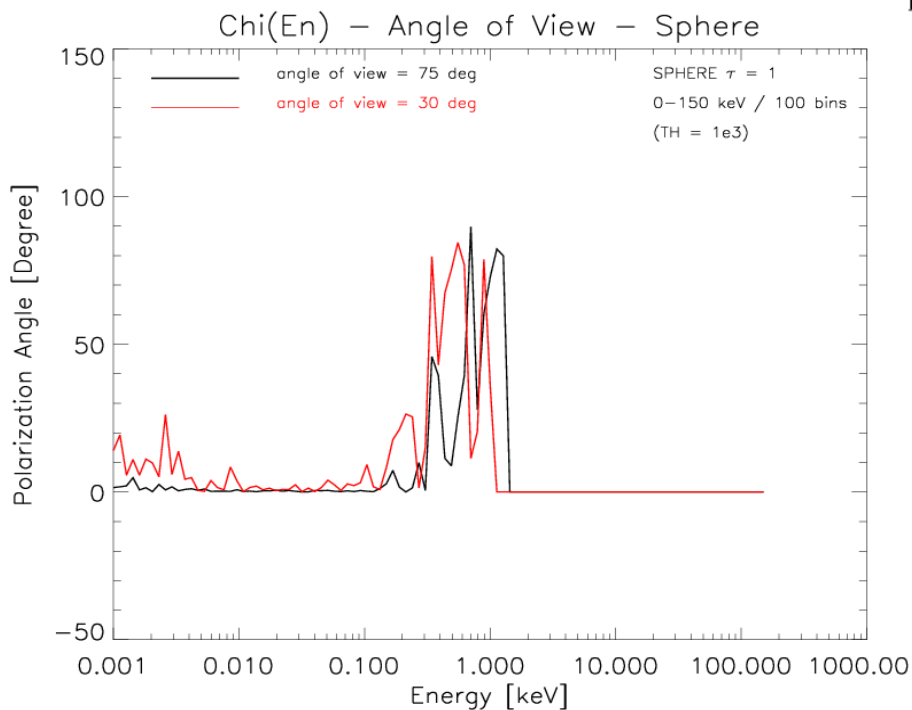
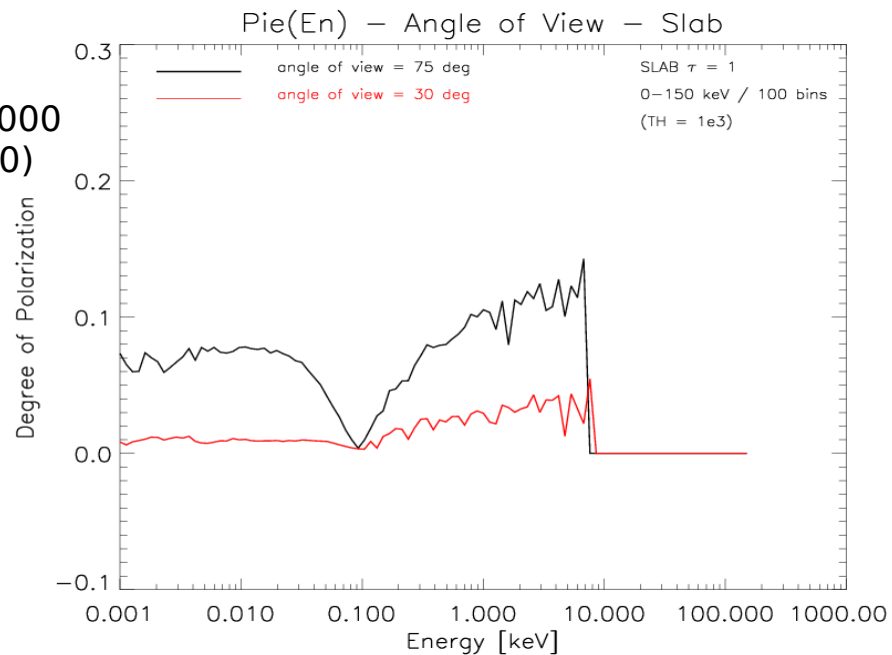
AGN Polarization

SLAB

$\tau = 1$

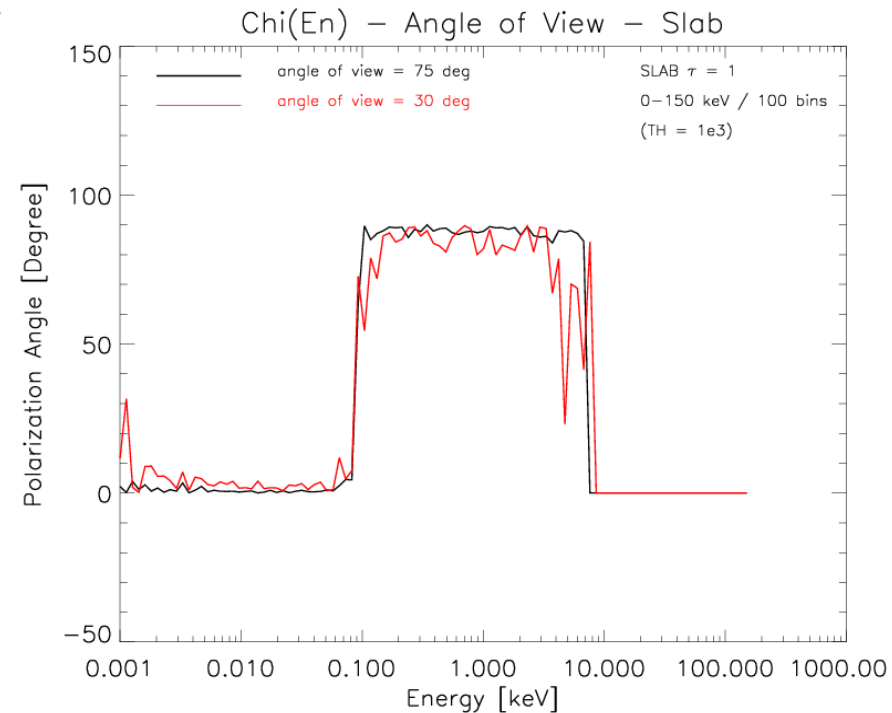


Threshold
(if counts < 1000
then set to 0)



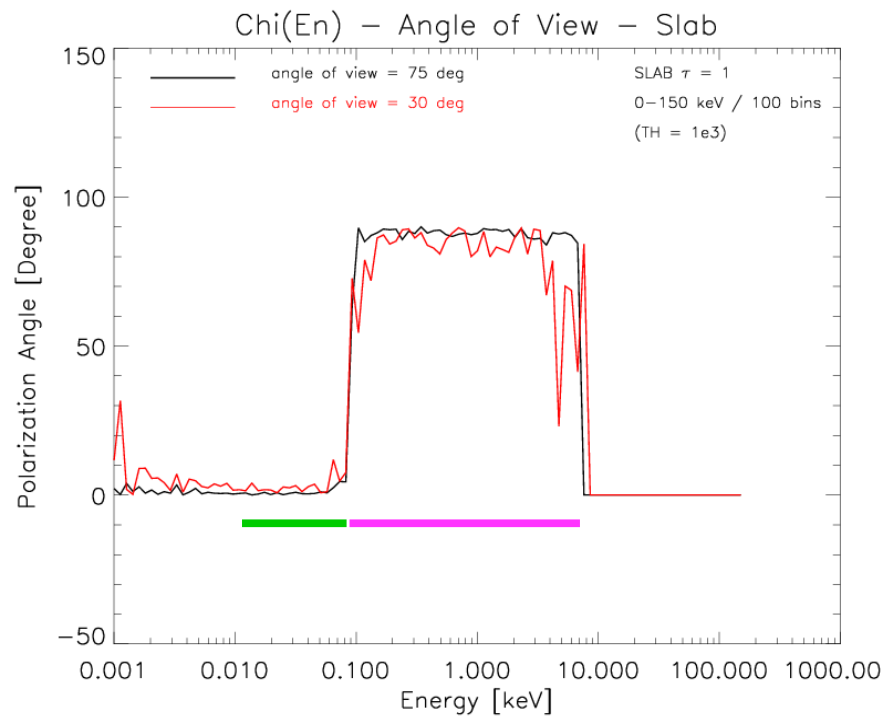
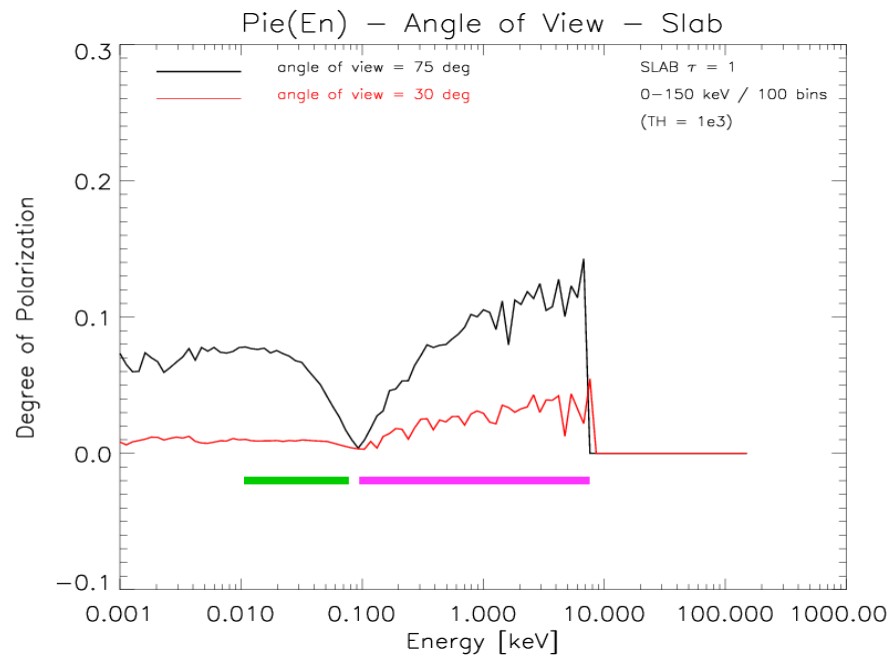
$$\Pi = \frac{\sqrt{Q^2 + U^2}}{I},$$

$$\chi = \frac{1}{2} \arctan \frac{U}{Q}.$$



AGN Polarization

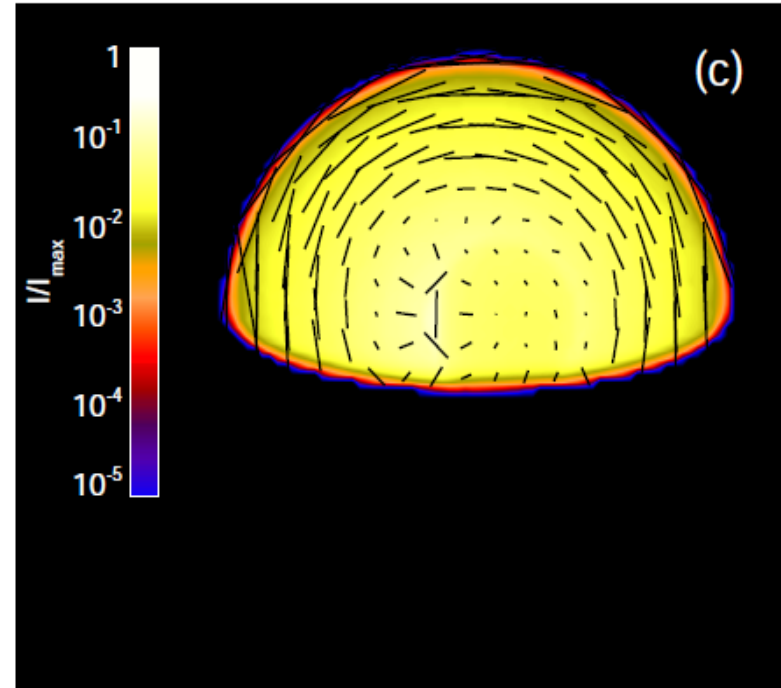
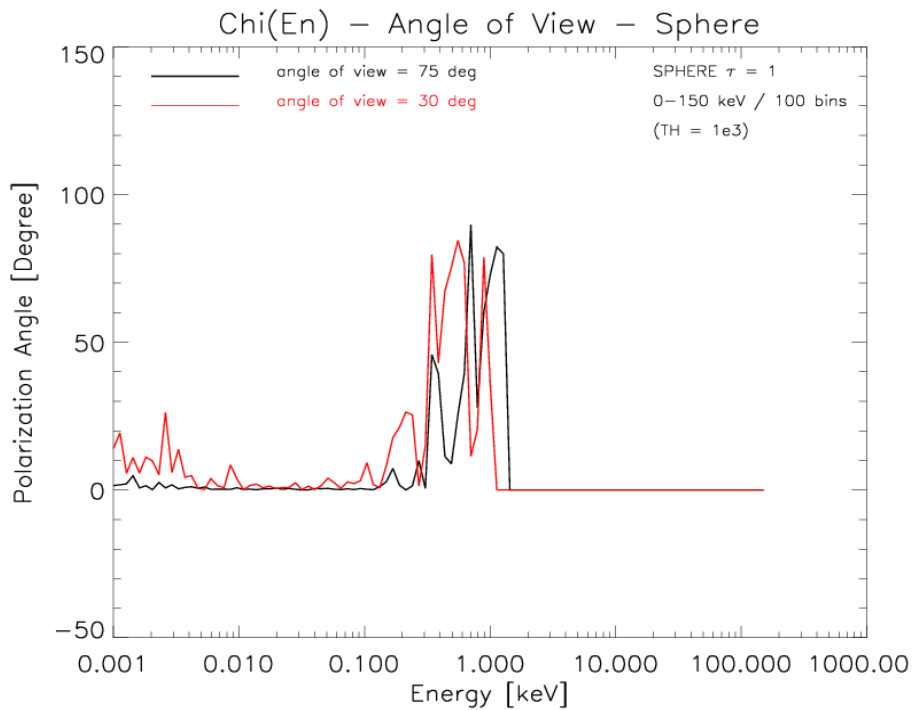
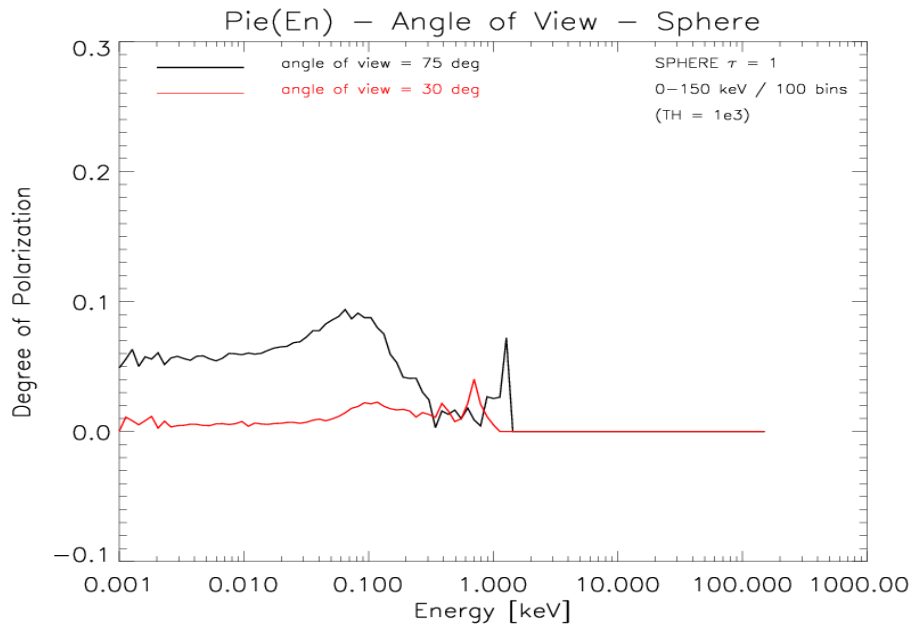
SLAB
 $\tau = 1$



SPHERE

$\tau = 1$

AGN Polarization



(picture from Schnittman+ 2009)

Iron line broadening in XRBs with NS

Relativistic Broadening or just Compton (down-)scattering?

Seed photons:

- monochromatic @ 6.4 keV
- unpolarized

Corona parameters:

- thermal energy ($kT = 2$ keV)
- optical depth ($\tau = 0.1, 1, 3$)

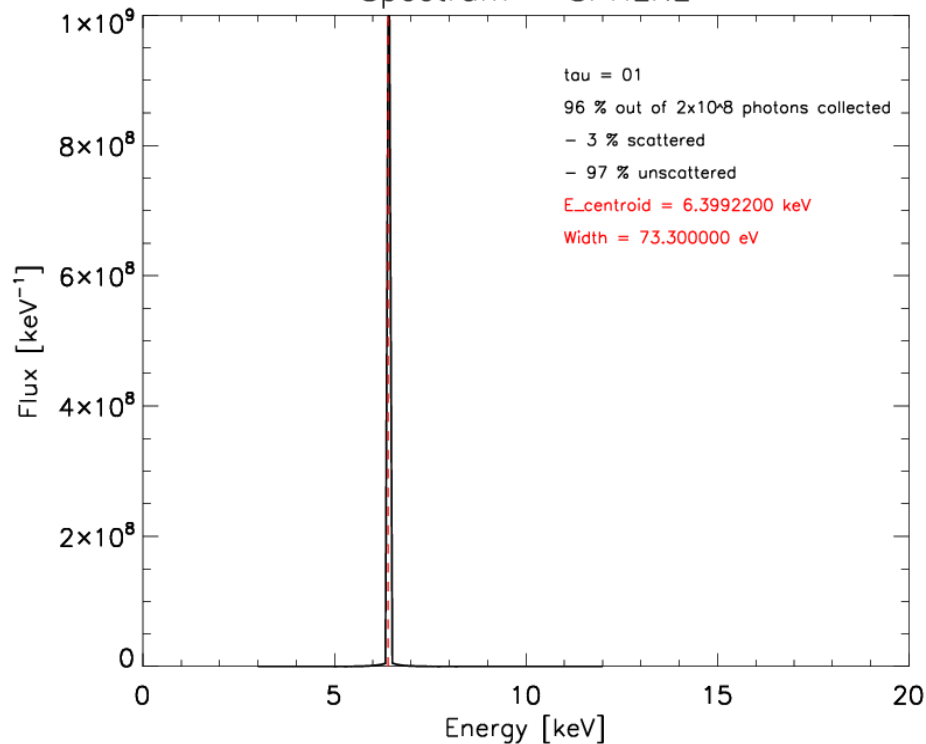
di Salvo+ 2009
D'Ai+ 2009
Cackett+ 2010
Egron+ 2011-13

?

Ng+ 2010

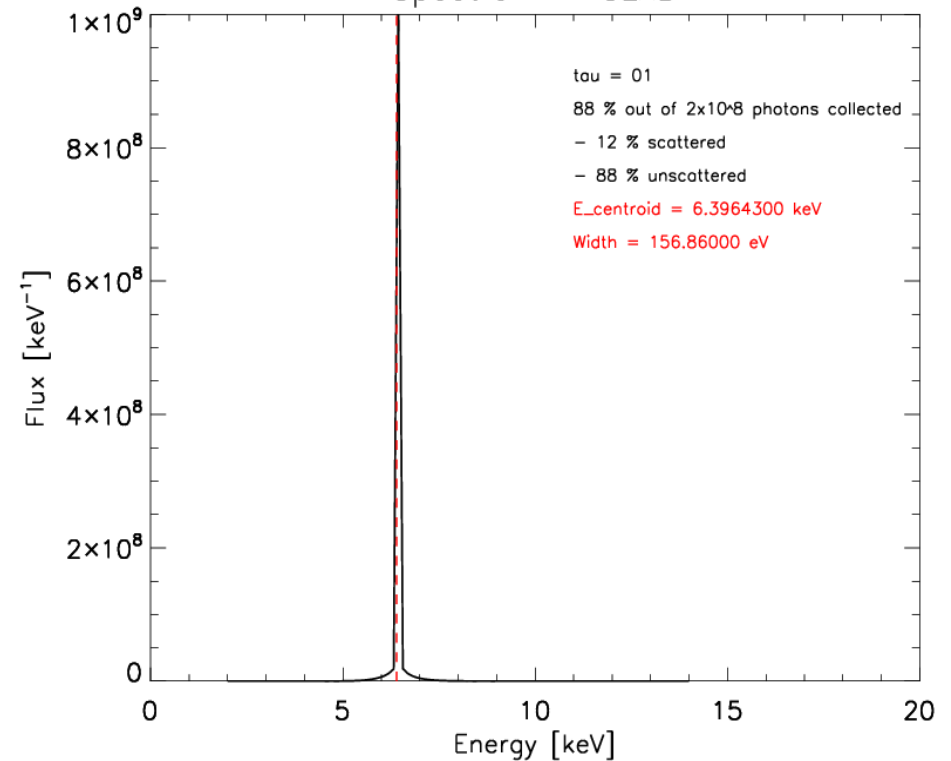
SPHERE
 $\tau = 0.1$

Spectrum - SPHERE



SLAB
 $\tau = 0.1$

Spectrum - SLAB

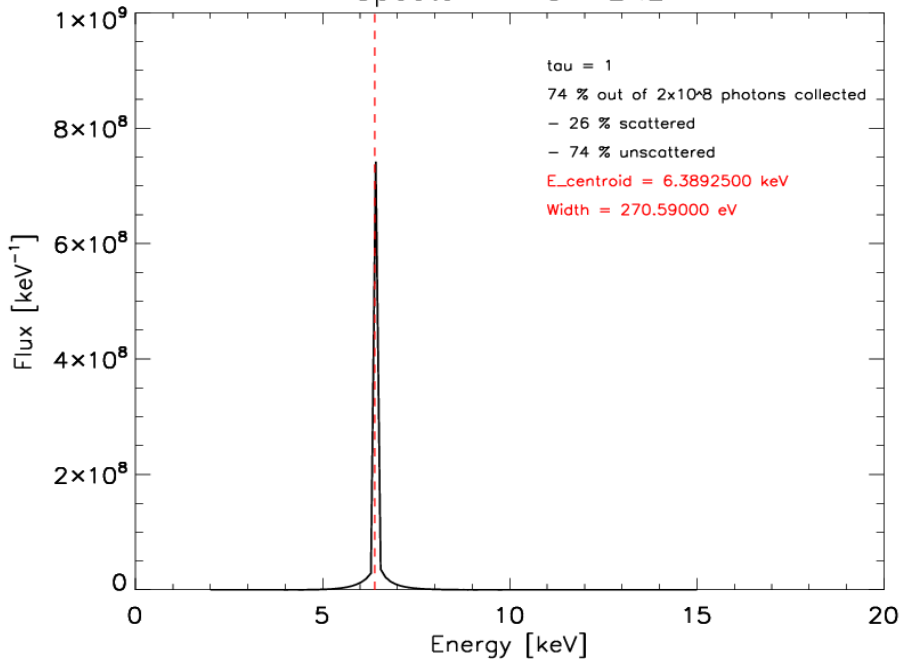


SPHERE

Line profile

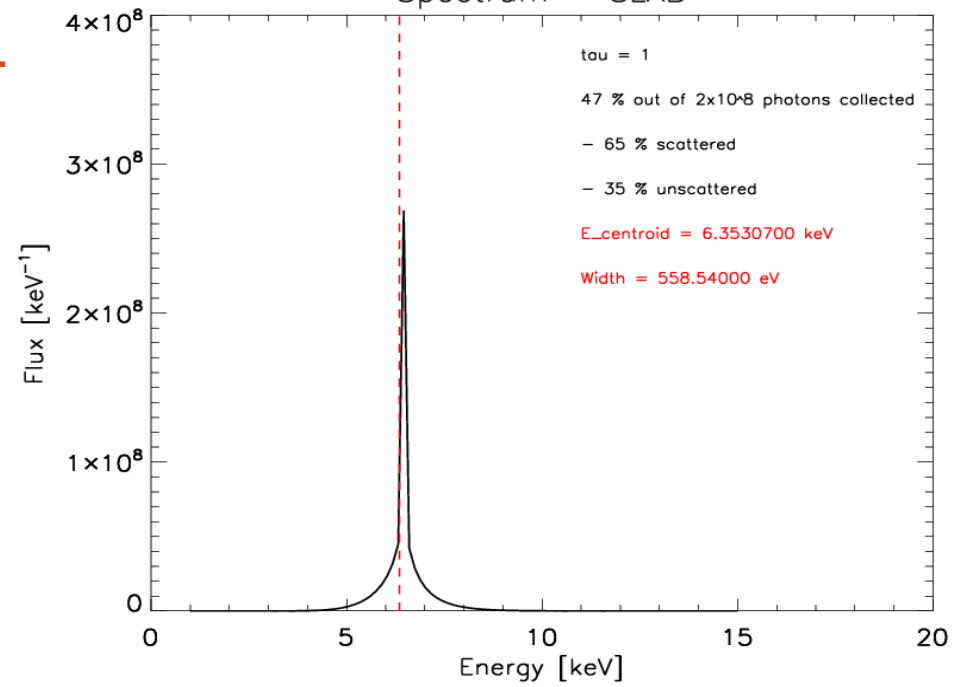
SLAB

Spectrum - SPHERE

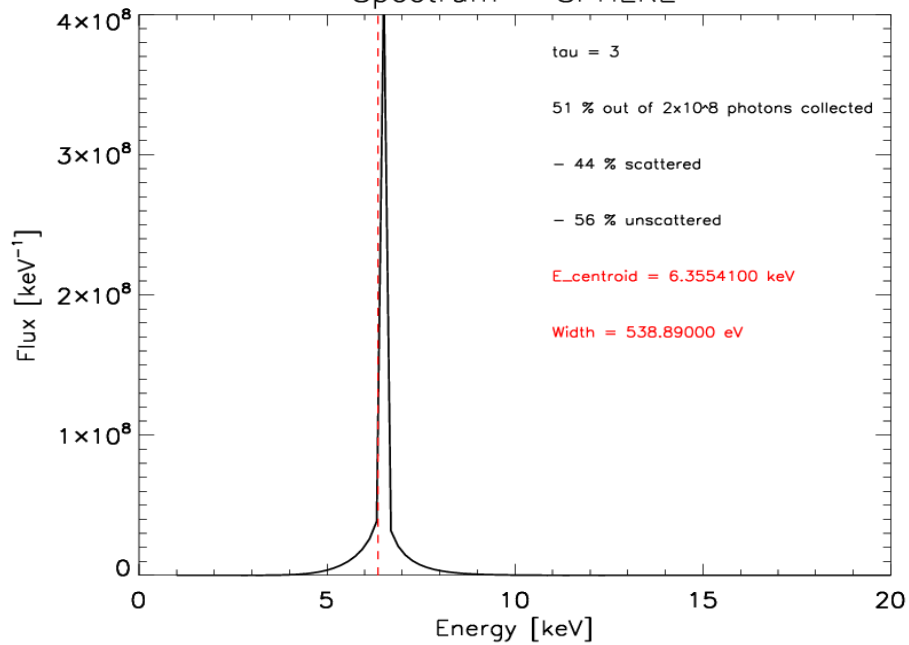


τ = 1

Spectrum - SLAB

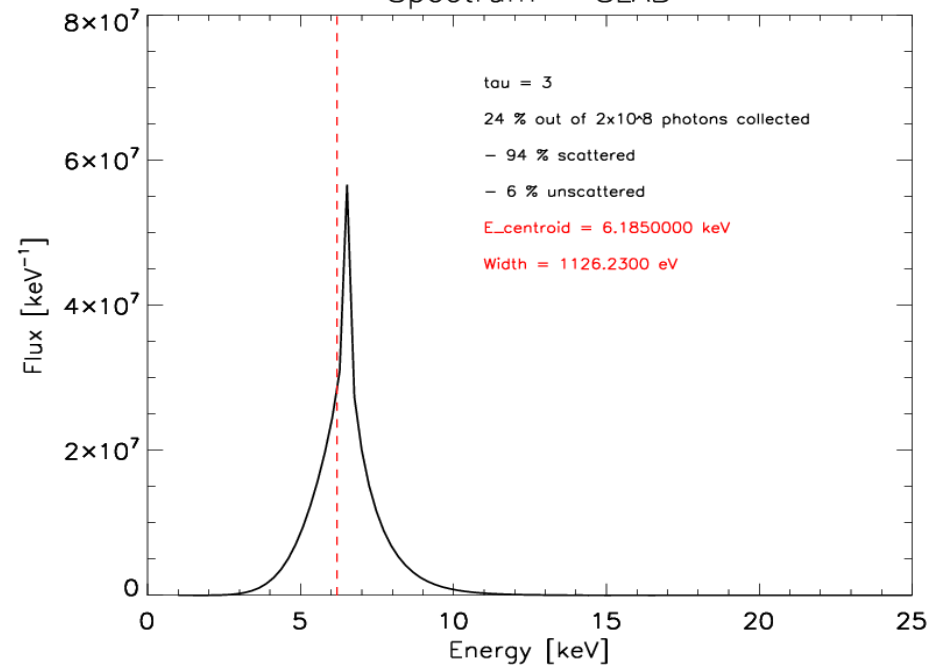


Spectrum - SPHERE



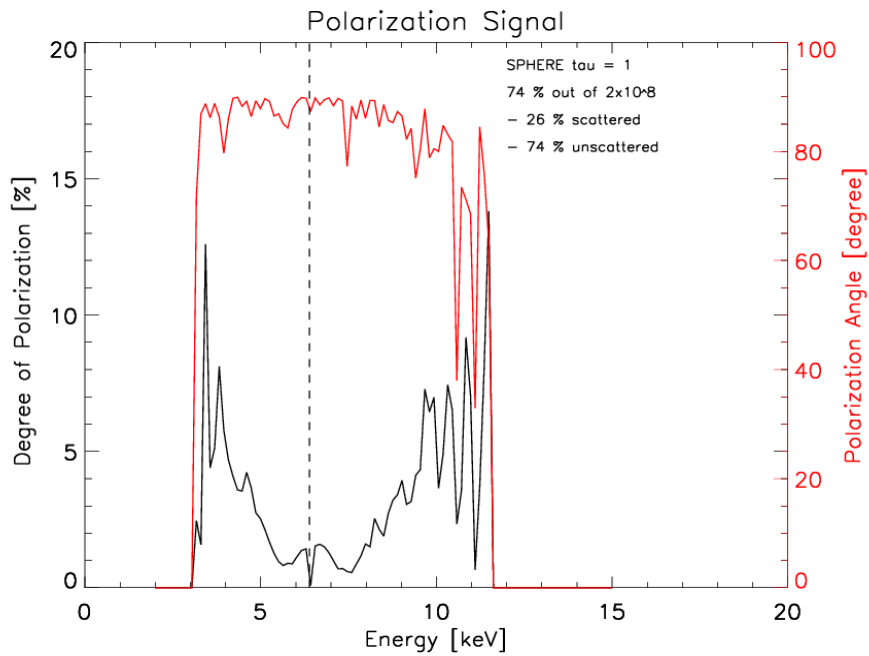
τ = 3

Spectrum - SLAB



Line polarization

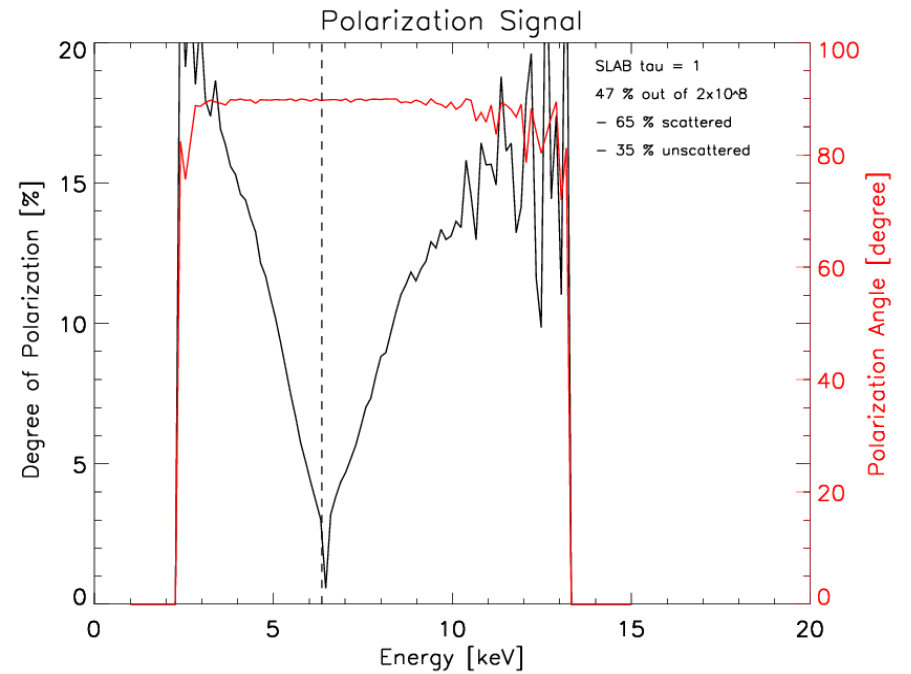
SPHERE



$\tau = 1$

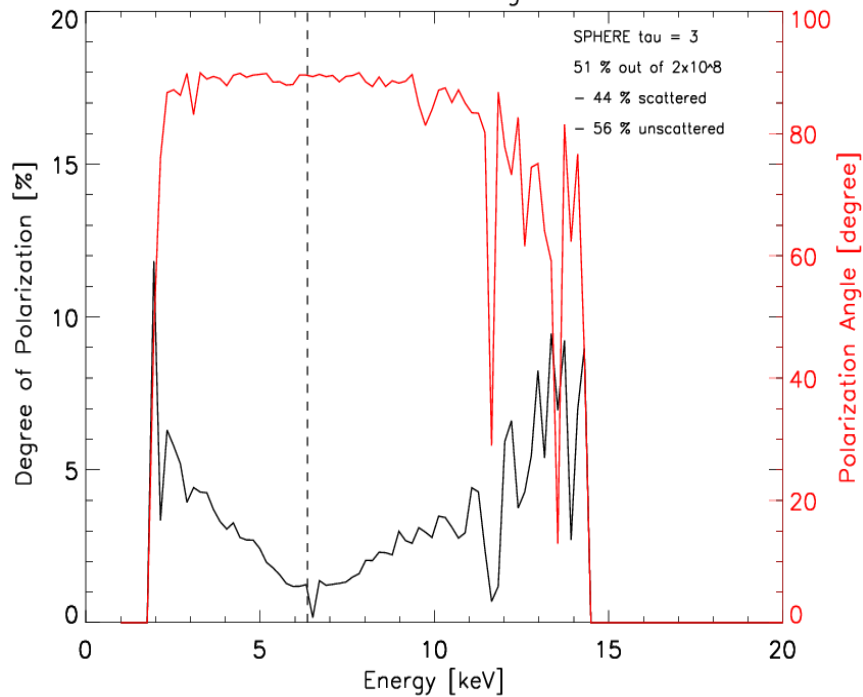
Threshold
100 counts

SLAB

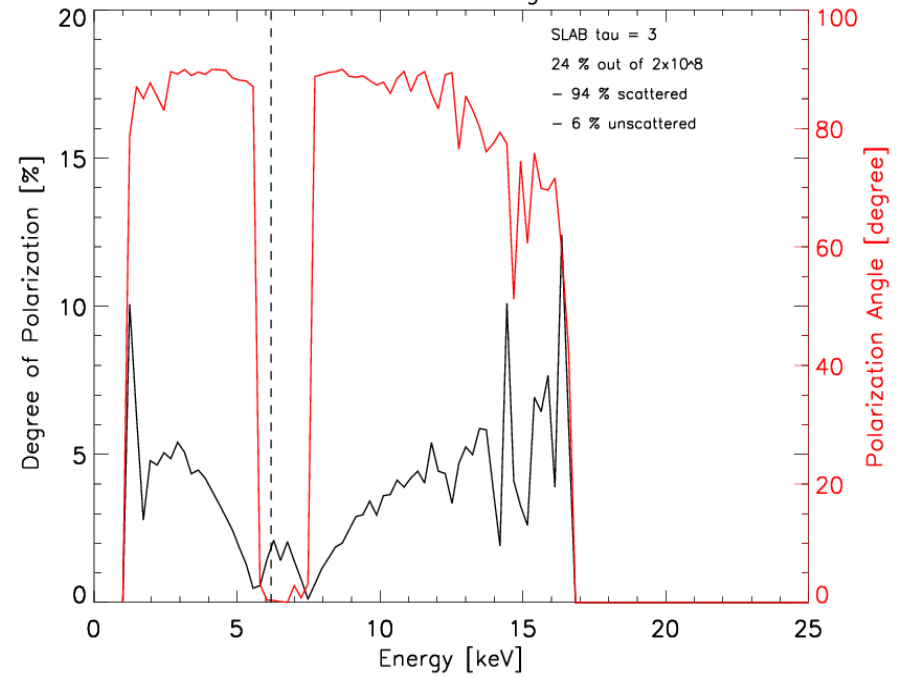


$\tau = 3$

Polarization Signal

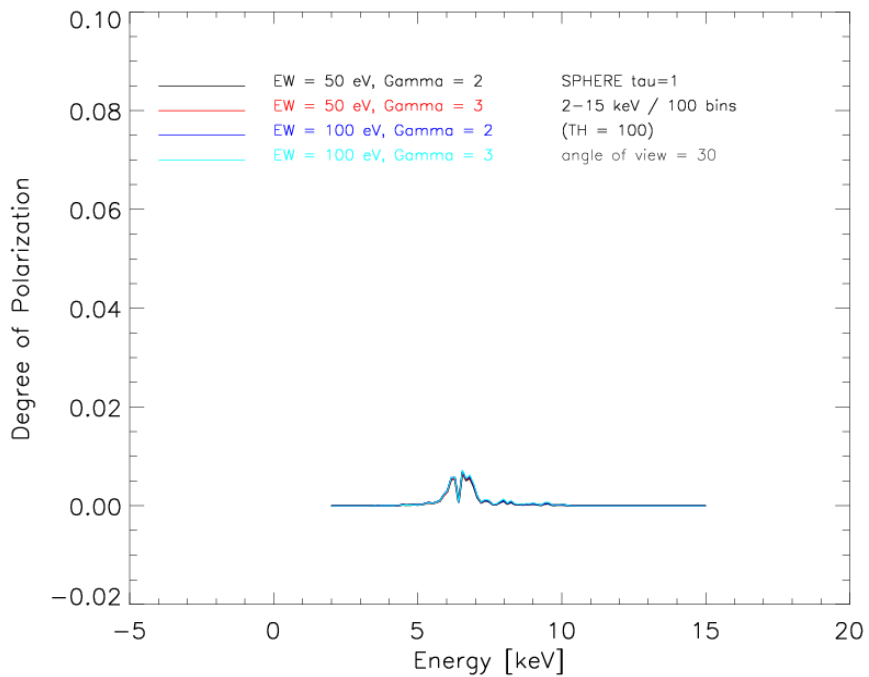


Polarization Signal

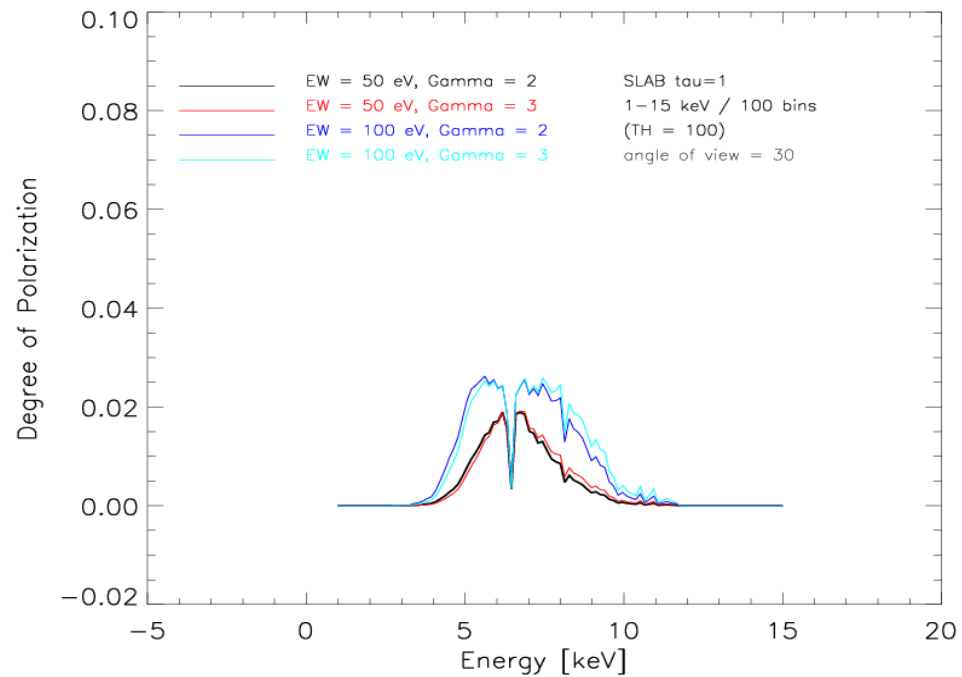


SPHERE**Net Polarization Degree - 30°****SLAB**

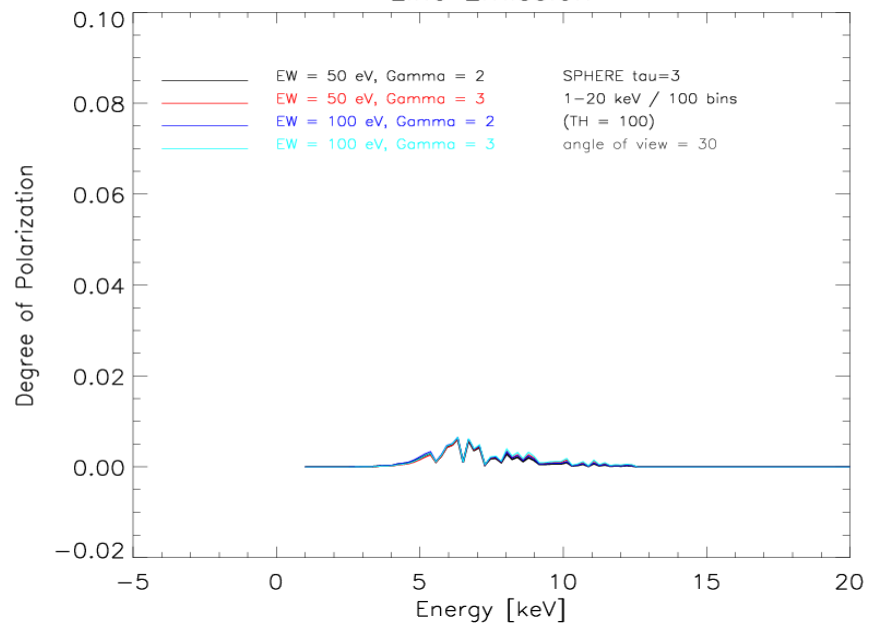
Line Emission

 **$\tau = 1$**

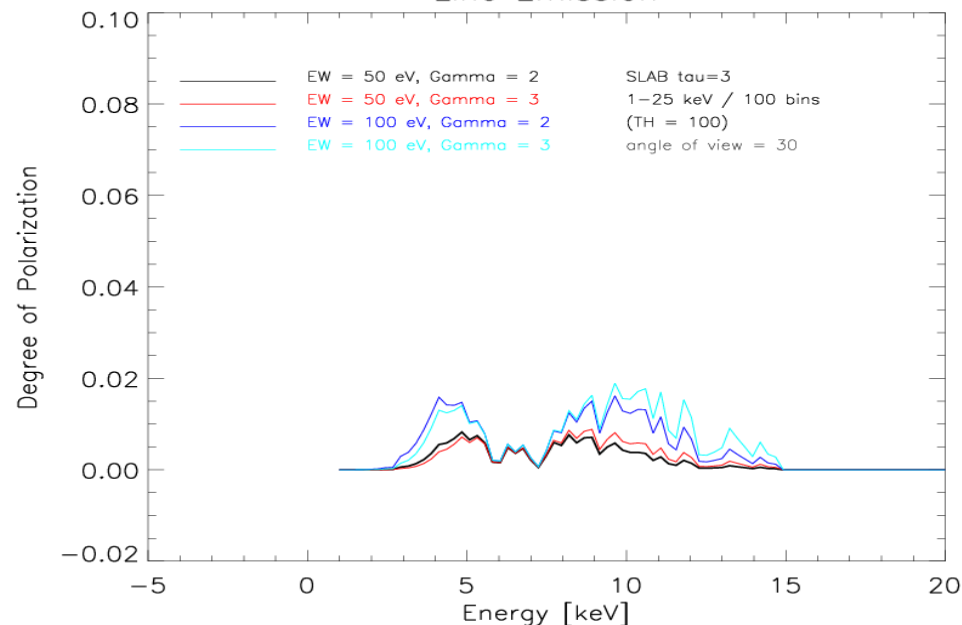
Line Emission



Line Emission

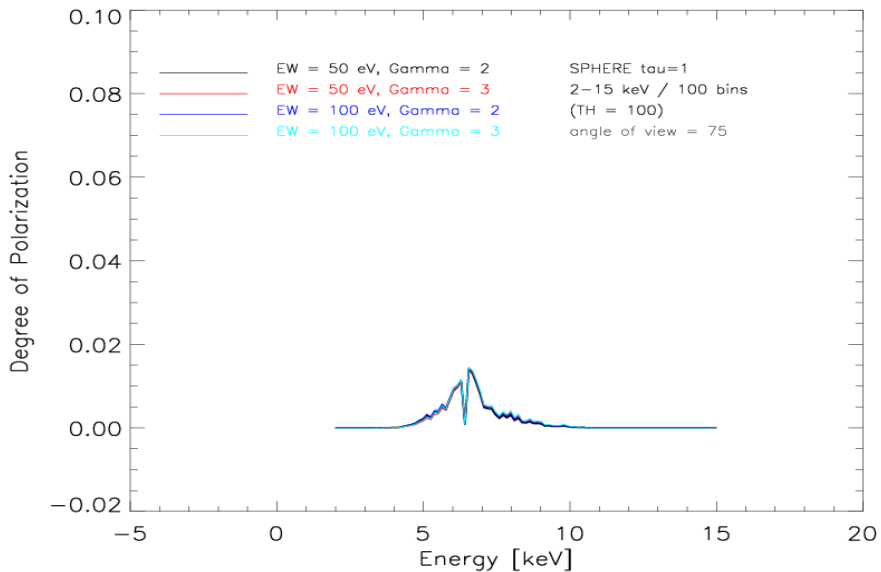
 **$\tau = 3$**

Line Emission

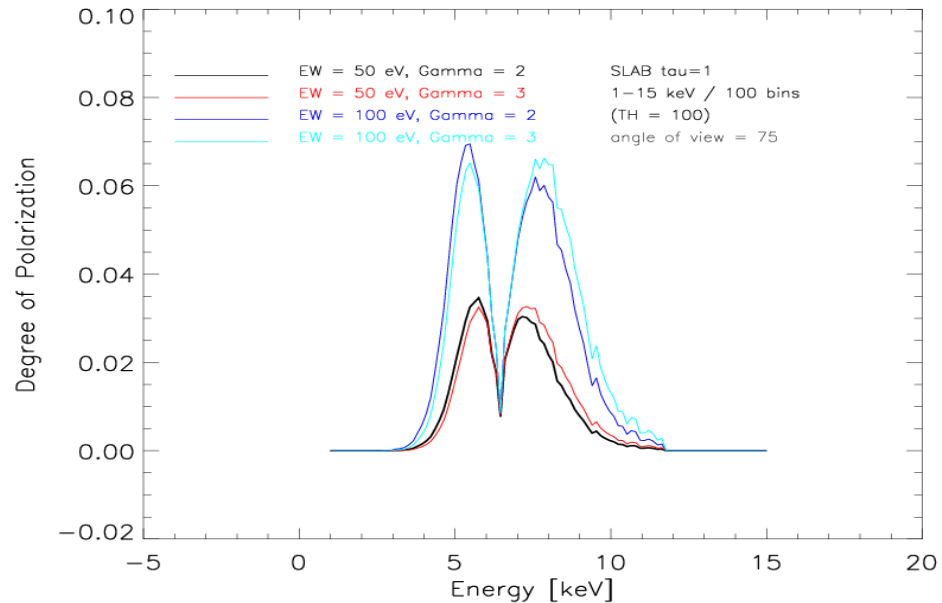


SPHERE**Net Polarization Degree - 75°****SLAB**

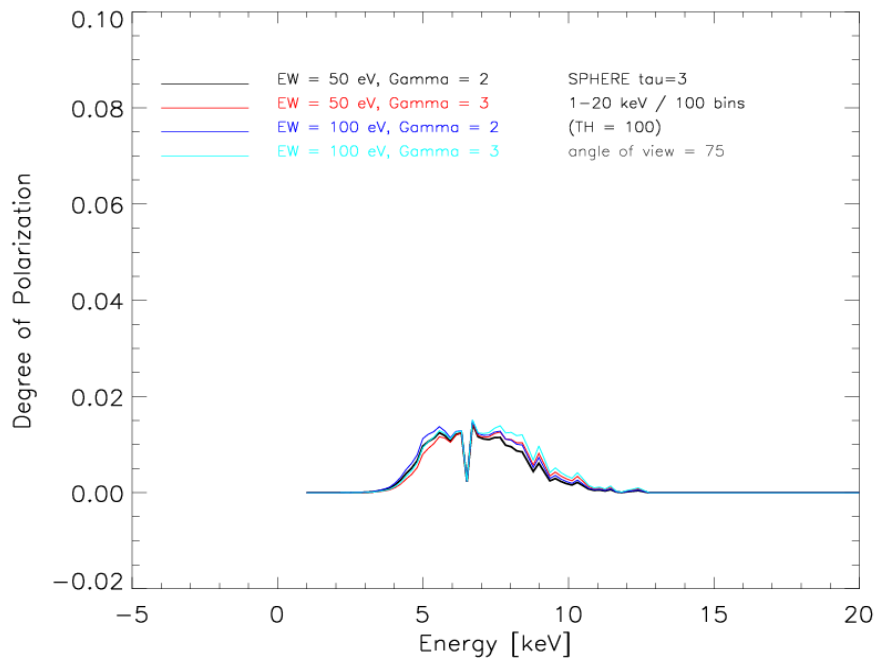
Line Emission

 **$\tau = 1$**

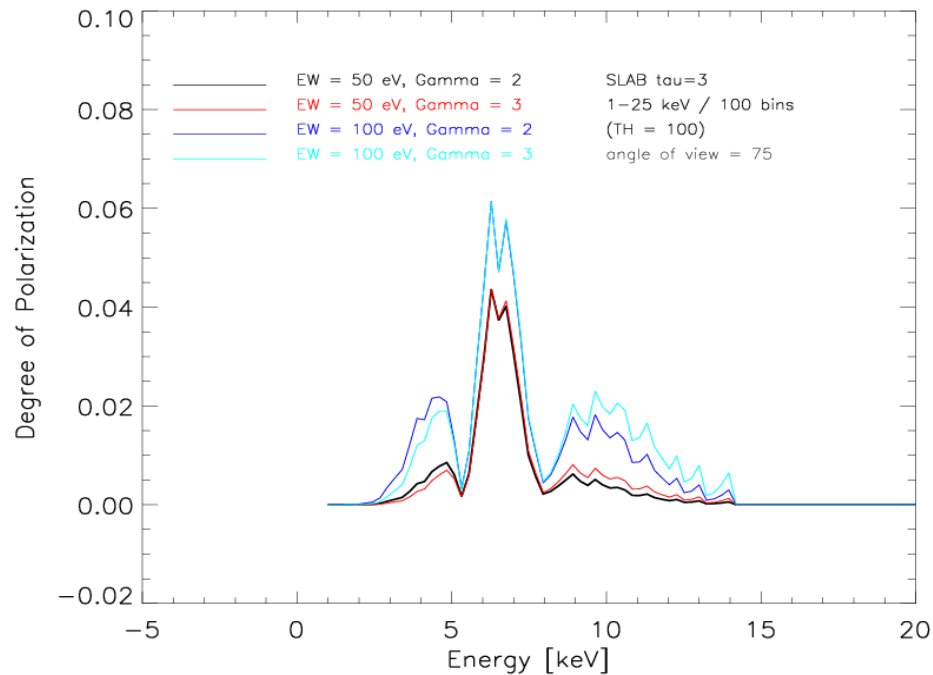
Line Emission



Line Emission

 **$\tau = 3$**

Line Emission



Conclusions

- Polarized radiation is expected in many sources being produced by scatterings, magnetic fields and general relativity
- Polarimetry offers 2 new independent observables (which is the opposite of adding more parameters!)
- These observables are extremely sensitive to geometry & angle of view if polarization is produced via scattering

Observational perspectives...none, yet!

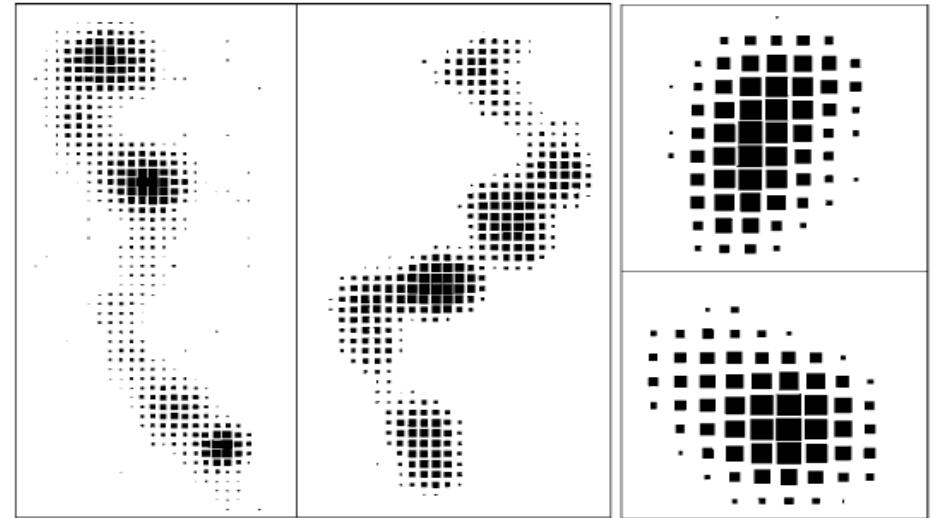
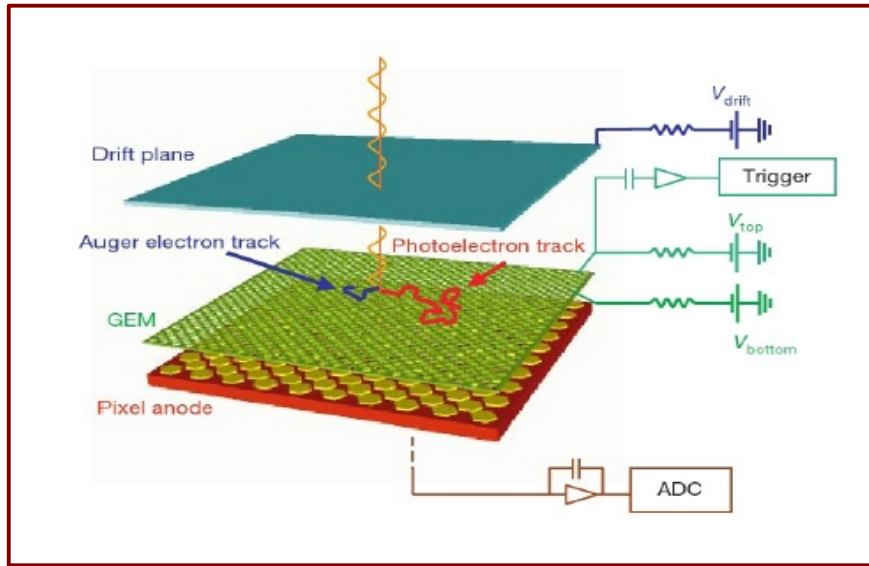


Fig. 3. Track images from 20 keV X-rays (left) and 4.5 keV X-rays (right).

Costa et al. (2001), Bellazzini et al.(2009), Muleri et al. (2009)

...but the technology is ready, well tested and perfectly fits a small mission (as we will see tomorrow afternoon).

Observational perspectives...none, yet!

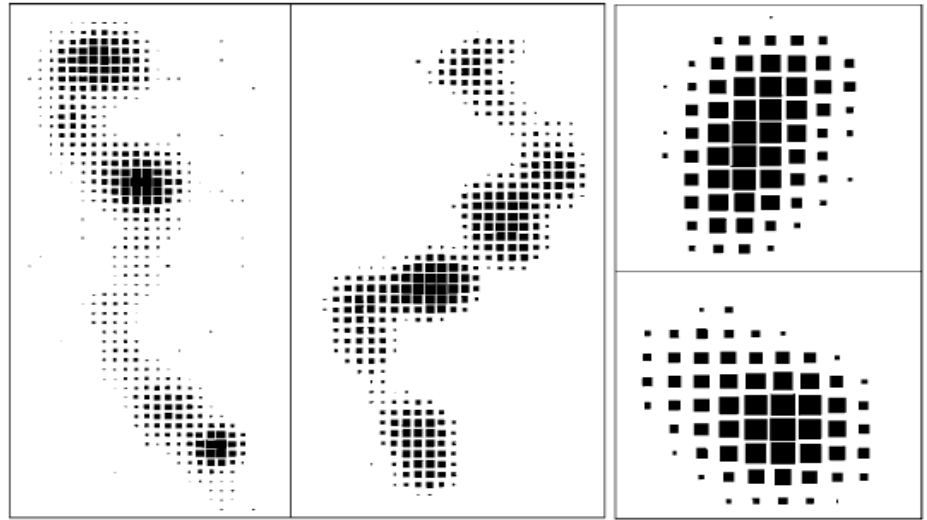
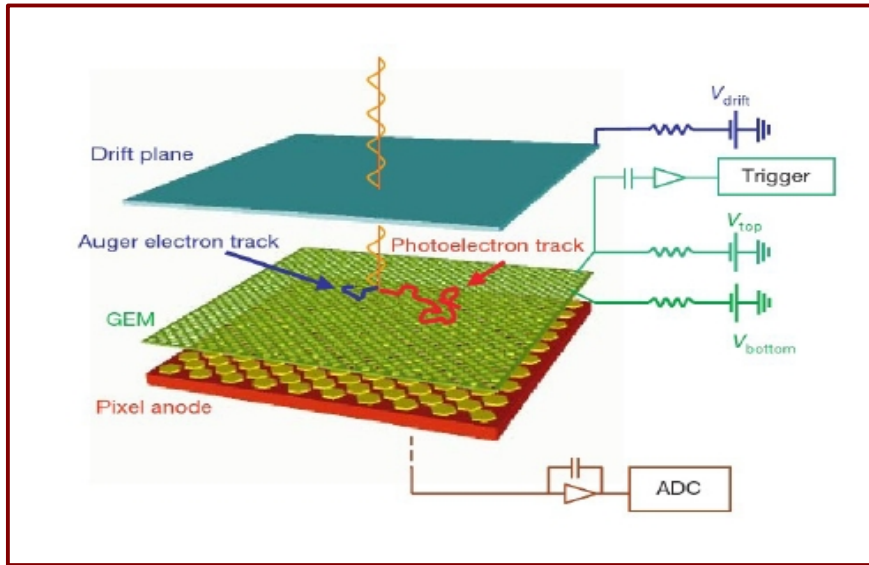


Fig. 3. Track images from 20 keV X-rays (left) and 4.5 keV X-rays (right).

Costa et al. (2001), Bellazzini et al.(2009), Muleri et al. (2009)

...but the technology is ready, well tested and perfectly fits a small mission.(as we will see tomorrow afternoon).

So let's open this window because:

“Even if the open windows of science at first make us shiver after the cozy indoor warmth of traditional humanizing myths, in the end the fresh air brings vigor, and the great spaces have a splendor of their own.”

[Bertrand Russell]